Reference orbit construction for tetrahedral formation on elliptical orbit

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Motivation

- 1) Slava G. Turyshev, Sheng-wey Chiow, and Nan Yu Searching for new physics in the solar system with tetrahedral spacecraft formations Phys. Rev. D, 109:084059, Apr 2024 doi: 10.1103/PhysRevD.109.084059
- 2) S.A. Bogachev, M.Yu. Ovchinnikov, V.A. Shuvalov Main Tasks and Possibilities to Study the Magnetic Field and Electric Currents in the Earth's Magnetosphere

Need for a tetrahedral formation

1) to study gravity gradient tensor e=0.6, a=1AU, r=1000km

2) to study Earth's magnetosphere e=0.7, a=20R_{Farth}, r=100km





by Small Spacecraft Cosmonautics and rocket science, 2024, 1(134), pp. 158-169 (in russian)

Tschauner–Hempel linearized equations

- elliptic reference orbit

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- in orbital frame

$$x'' - 2y' - \frac{3x}{1 + e \cos \nu} = 0$$
$$y'' + 2x' = 0$$
$$z'' + z = 0$$

- bounded solution for satellite number j has the form

$$x_j(\nu) = A_j(1 + e\cos\nu)\sin\nu + B_j(1 + e\cos\nu)\cos\nu$$
$$y_j(\nu) = A_j(2 + e\cos\nu)\cos\nu - B_j(2 + e\cos\nu)\sin\nu + C_j$$
$$z_j(\nu) = D_j\sin\nu + E_j\cos\nu$$

Need to maintain constant tetrahedral volume

to make accurate measurements of gravitational/electromagnetic field

We did such a thing for circular orbits: S. Shestakov, M. Ovchinnikov, Y. Mashtakov Analytical Approach to Construction of Tetrahedral Satellite Formation Journal of Guidance, Control and Dynamics, 2019, V. 42, No. 12, pp. 2600-2614 doi: 10.2514/1.G003913

Every satellite with j=1,2,3 has its own A_j, B_j, C_j, D_j, E_j Combine constants into vectors $\mathbf{A} = (A_1, A_2, A_3)^T$

Then for the volume we have

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\det \|\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3\| = (1 + e \cos \nu) \cdot (
                                                                                                \frac{1}{2} \cdot (\langle \mathbf{A}, \mathbf{C}, \mathbf{D} \rangle + \langle \mathbf{B}, \mathbf{C}, \mathbf{E} \rangle - e \langle \mathbf{A}, \mathbf{B}, \mathbf{E} \rangle) +
         We calculate
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We decompose a numerator and a denomenator into trigonometric polynomial of degrees of true anomaly

We get the necessary and sufficient conditions for the volume to conserve, k is arbitrary here

So, distance between satellites to orbit semi-axis ratio is a small parameter



the volume,	
here angle brackets	
are mixed product	
of vectors	
are mixed product of vectors	

$-2\cos\nu\cdot\langle \mathbf{A} \mathbf{B} \mathbf{E} \rangle +$	or degrees of the anomaly
$-2\sin\nu \cdot \langle \mathbf{A}, \mathbf{B}, \mathbf{D} \rangle +$	$\langle {f A}, {f B}, {f D} angle$
$\frac{1}{1} \qquad ((A \cap D) + (D \cap D) + (A \cap D)) + (A \cap D) + ($	$\langle {f A}, {f C}, {f E} angle$
$\frac{1}{2}\cos 2\nu \cdot (-\langle \mathbf{A}, \mathbf{C}, \mathbf{D} \rangle + \langle \mathbf{B}, \mathbf{C}, \mathbf{E} \rangle - e \langle \mathbf{A}, \mathbf{B}, \mathbf{E} \rangle) +$	$\langle {f A}, {f B}, {f E} angle$
$\frac{1}{2}\sin 2\nu \cdot (\langle \mathbf{A}, \mathbf{C}, \mathbf{E} \rangle + \langle \mathbf{B}, \mathbf{C}, \mathbf{D} \rangle - e \langle \mathbf{A}, \mathbf{B}, \mathbf{D} \rangle))$	$\langle {f A}, {f C}, {f D} angle$

= 0 $+\langle \mathbf{B}, \mathbf{C}, \mathbf{D} \rangle = 0$ =-ke= k $\langle \mathbf{B}, \mathbf{C}, \mathbf{E} \rangle = k$

Numerical experiment

Motion of tetrahedral formation around the Sun a=1AU, e=0.6, central gravity initial conditions are such that volume is conserved in Tschauner-Hempel model

In nonlinear model the volume is still constant



Not only volume must be maintained

Formation shape is also of interest in applications e.g. for intersatellite links to work distances and relative angles must be under control

But shape is not a single number We choose quality parameter

$$\mathbb{Q} \sim \mathbb{V}^{2/3} \cdot \mathbb{L}_{\mathrm{sq}}^{-1}$$

Here L_{sq} is sum of squares of lengths of all edges

With constant volume it is impossible to maintain constant quality unless the reference orbit is circular

But it is possible to analytically minimize the variations in quality



400 t, days Difference between numerical and theoretical values of L_{sa}

